0.9 MILES BELOW 1.5 KM THE 1 KM <sup>2</sup> 1.5 KM IDE FOOTPRINT 86 CABLES LIGHT SENSORS		7,500 atmospheric Vabout 3 0.9 MILES BELOW NUONS/MONTH 7,500 atmospheric Vabout 3 0.9 MILES BELOW NEUTRINOS/MONTH 1.5 KM THE
1.5 MILES BELOW 2.5 KM THE ICE SOUTH POLE NEUTRING DESERVATORY	EVENTS LAST ABOUT D.DODOO3 SECONDS	ICECUBE.WISC.EDU



FRONT VIEW

BACK VIEW



## ICECUBE



Cosmic neutrinos can produce different patterns of light in IceCube. Event displays tell us about important properties of these neutrinos.

## TRACK

When a muon neutrino interacts with the ice, it creates a muon as a secondary particle. The muon crosses the detector leaving a track of light in IceCube.



This neutrino was detected in IceCube on Friday, November 12, 2010



This is a simulated tau neutrino.

This neutrino was detected in IceCube on Tuesday, December 4, 2012.

## DOUBLE BANG

A tau neutrino produces a cascade pattern with two large energy deposits. A double-bang pattern would be an unambiguous signature of a tau neutrino but it has not yet been observed in lceCube.



## DIGITAL OPTICAL MODULE,

IceCube sensors are the size of a basketball. They work as an inverse light bulb that turns the blue shimmering light produced by a neutrino interacting with the ice into a digital signal.

Kilometer-long cables connect DOMs to the IceCube Lab, where data is prepared for satellite transmission to iceCube headquarters at the University of Wisconsin-Madison.



