# High Energy Neutrinos in IceCube: How can we find them?

Masterclasses 2015 Nancy Wandkowsky

### What do we know about neutrinos?



1912: Hess discovers "Cosmic Rays" (mainly p)





Where do they come from and how are they accelerated?

Ingredients for cosmic neutrinos:







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- cosmic proton accelerator
- nearby target, 'beam dump'

 $p + \gamma \rightarrow \pi^+ + n \rightarrow \mu^+ + \nu_{\mu} + n$ 





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# Neutrinos are ideal messenger particles!





neutrinos themselves leave no trace → 10<sup>20</sup> neutrinos pass through IceCube every second, undetected!

so what is IceCube measuring?



first, the neutrino has to interact with the ice, producing charged particles such as muons

muons pass us at a rate of 1 per cm<sup>2</sup> and minute: why don't we see them, but the detector can?



# high-energy charged particles traveling in a **medium** produce Cherenkov light!



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### What do neutrinos look like in IceCube?

track

 $\nu_{\mu} + N \rightarrow \mu + X$ 





 $v_e + N \rightarrow e + X$  $v_x + N \rightarrow v_x + N$ 

x = any flavor

## What is IceCube seeing?

 For each event, we want to know the energy and direction: how good does this work for tracks/ cascades?





good direction resolution bad energy resolution

# good energy resolution bad direction resolution

### What do neutrinos look like in IceCube?

track





- each coloured dot represents a sensor detecting light
- size scales with the amount of recorded light
- color indicates arrival time: red first, green last

# now it's your turn to identify events in this IceCube event display

### http://icecube.wisc.edu/viewer/quiz (Particle ID Quiz)



# now it's your turn to identify events in this IceCube event display

#### http://icecube.wisc.edu/viewer/quiz

# Are all of the events you just saw astrophysical neutrinos?





signal: astrophysical v; background: everything else

Background:

- atmospheric neutrinos
- atmospheric muons (many! can get two at the same time: coincident event)



#### challenge: select only signal events!



How can we distinguish signal from background?

#### can you guess what is signal and what is background?

http://icecube.wisc.edu/viewer/background\_signal

(Background vs Signal)



Finding needles in a haystack



275 million atmospheric muons are detected daily, created by interactions of cosmic rays with the earth's atmosphere 8,250 atmospheric neutrinos are detected monthly



only 10s of cosmic neutrinos are detected per year

Any ideas how to select the signal events?

#### typical background



#### typical signal







Idea: events starting in the detector have to be neutrinos





#### use DOMs on outer layer as veto!





an event is marked as signal if:

- less than 3 of the first 250 photoelectrons were recorded in the veto region
  - → removes atmospheric muon background



an event is marked as signal if:

- less than 3 of the first 250 photoelectrons were recorded in the veto region
  - → removes atmospheric muon background
- it deposits more than 6000 PE
  - → removes some atmospheric neutrino background



#### now it's your turn: which events pass/fail the veto?

#### http://icecube.wisc.edu/viewer/training

(Selecting neutrinos)



when opening up the box, we found 28 events in 2 years of measurement

have a closer look at the individual events:

http://icecube.wisc.edu/viewer/hese

(28 very high energy events)



when opening up the box, we found 28 events in 2 years of measurement

we are interested in the following properties:

- energy
- time of event (MJD)
- arrival direction (RA, Dec)
- angular error
- event type

MJD: days since 0h Nov 17, 1858 (now: ~57099.6)



when opening up the box, we found 28 events in 2 years of measurement

can you guess which ones have the highest energy?

http://icecube.wisc.edu/viewer/hese\_all

(The highest energy events)



when opening up the box, we found 28 events in 2 years of measurement

can you guess which ones have the highest energy?

#### http://icecube.wisc.edu/viewer/hese\_all

why is it not easy to identify the highest energy events by eye?

- ice is not the same everywhere
- energy can be deposited outside the detector

• . . .

we call these "systematic effects"

"How significant is our result?"

How much background did we expect? → needs precise computer models



"How significant is our result?"

How many events did we actually find?

→ can this be explained by a background-only hypothesis?



"How significant is our result?"

To achieve the best results, we have to tune the event selection:

http://icecube.wisc.edu/masterclass/tuning (Tuning the event selection)

"How significant is our result?"



#### deposited energy (GeV)



"How significant is our result?"

Can you think of another tuning parameter?

http://icecube.wisc.edu/masterclass/tuning\_zen

Science 2013 510

Yes, we did!



### The IceCube Neutrino Observatory



## The IceCube Neutrino Observatory

- 5160 light detectors
- 1 km<sup>3</sup> volume
- 86 strings
- 125 m string spacing
- completed in 2010



### What do neutrinos look like in IceCube?



#### muons: long paths in the detector $\rightarrow$ track

### What do neutrinos look like in IceCube?



#### electrons/hadrons: shower of light $\rightarrow$ **cascade**