Status of the IceTop air shower array at the South Pole

Fabian Kislat, DESY 13th ICATPP, Como October 3rd, 2011





IceCube and IceTop

- 1km² air shower array
- 1km³ neutrino telescope
- 3D air shower detector
- 2835 m.a.s.l. 690g/cm²
- 125m detector spacing
- Energy range 300TeV 1EeV



TeV muon bundle in IceCube



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IceTop Array



IceTop Energy Range

- Threshold: 300TeV
 Infill: ~100TeV
 - Overlap with direct
 measurements
- Max. energy: O(1EeV) Limited by statistics
 - Connection to Auger
- Knee and 2nd knee
- End of galactic cosmic rays?
- Spectrum and composition: Important information on acceleration mechanisms



IceTop Tanks



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Signals and Calibration



- DOMs record 422ns waveforms
- Signals dominated by photons
- Calibration using single muons
 - Large number of muons from low energy primaries
 - Distinctive light yield from muons
 - Position of "muon peak" = 1 Vertical Equivalent Muon

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AIR SHOWER RECONSTRUCTION







Air Shower Footprint



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Lateral Distribution Function



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Shower Front



• Fixed shape:

$$\Delta t(R) = a R^2 + b \left(\exp\left(-\frac{R}{2\sigma^2}\right) - 1 \right)$$

• $a, b, and \sigma$ are constants

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Reconstruction Performance

26-station configuration



- After containment cuts:
 - Core position resolution < 10m
 - Angular resolution < 1°

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SPECTRUM & COMPOSITION





Composition Measurement with IceTop



GeV muons /



Inclined showers



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IceTop-IceCube Coincidences



- Neural network
- Inputs:
 - Shower size S₁₂₅
 - Muon bundle size K70

- Outputs:
 - Primary energy
 - Primary mass

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IceTop-IceCube Coincidences



- 40-Station configuration
- Sub-Array with only small amount of snow on tanks
- 1 month of data
- Systematics dominated

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Muon Counting in IceTop



- GeV muons in IceTop tanks
- Muons
 ⇔ Large signals at shower periphery
- 1 VEM charge enhanced if $\langle Signal \rangle_{em} \ll \langle Signal \rangle_{\mu}$
- Statistical muon counting
- Will profit from **SLC**: Charge stamps if no Local Coincidence

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Inclined Showers



- Energy spectra in different zenith angle ranges
- Heavier primaries
 ⇒ stronger shower attenuation
- Isotropy
 - \Rightarrow spectra must agree
- Composition sensitivity from shower development
- IceTop-26 analysis 1-100PeV

OTHER ANALYSES







$High-p_T$ muons



- Muon separation spectrum > 150m
- Test of models
- High-p_T muons from charm decays
- IceTop delivered direction reconstruction

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Low-energy transient events



- Discriminator rates sensitive to GeV primary flux
- Additional discriminators:
 - Various thresholds
 - Spectroscopy
- Solar flares
- Supernovae
- GRBs

Veto for neutrino search



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- Earth not transparent for PeV neutrinos
- Look above horizon
- Large atmospheric muon background
- Use IceTop as veto
- At sufficiently high energy:
 - ~ 100% background rejection
 - > 90% signal efficiency
- SLC: Improve capabilities
 - Larger zenith angles

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PeV Photon search with IceCube veto





- Photon showers are muon poor
- Use IceCube as a veto
- Upper limit on E⁻² flux at 1.2PeV (99% c.l.):
 Φ < 7.0 · 10⁻¹⁹ cm⁻² s⁻¹ TeV⁻¹

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Summary

- IceCube/IceTop construction complete
- First results:
 - Cosmic-ray energy spectrum and composition
 - Physics of air showers
 - Heliospheric physics, transient events
 - EAS veto for high-energy neutrino searches

Ongoing efforts:

- Composition measurement with muons in IceCube
- Cosmic-ray anisotropy measurement
- Future plans:
 - Horizontal air showers
 - Radio extension (RASTA)

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