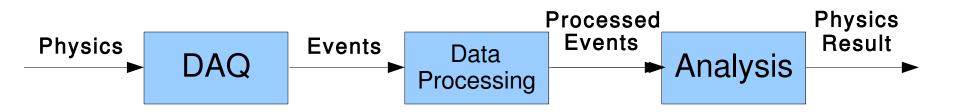
Data Challenge

Jim Braun
2009 IceCube Science Advisory Committee Meeting

Data Challenge

- Most IceCube physics analyses lack a natural signal
- Data challenge is the following generic process:
 - A simulated signal is added to the data
 - The data is analyzed, and the signal is either discovered at the expected level, or it is not
- Data challenge accomplishes the following:
 - Detects errors in DAQ, data processing, and analysis
 - Ensures those performing analysis understand their data and methods before physics results are obtained
 - Aids interpretation of physics results
 - Unbiased measurement of the performance of an analysis

What Can Go Wrong?

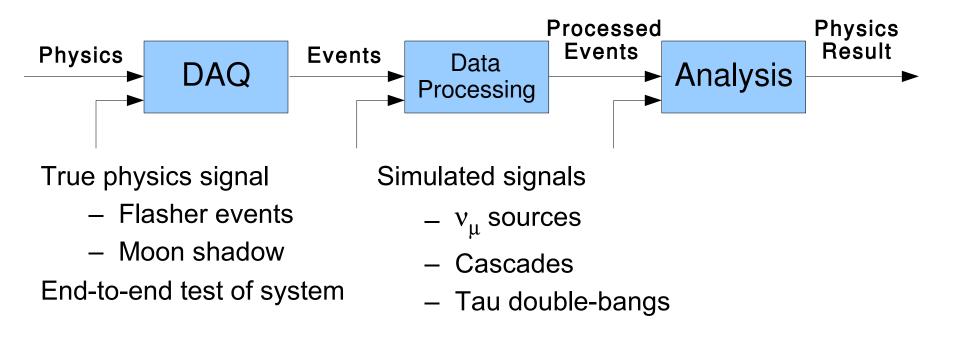


- System timing
- DAQ hardware and software failures
- Angular reconstruction
- Energy reconstruction
- Coordinate transformations

- Data selection
- Analysis methods

 We want to ensure the entire system is working properly from physics to analysis result

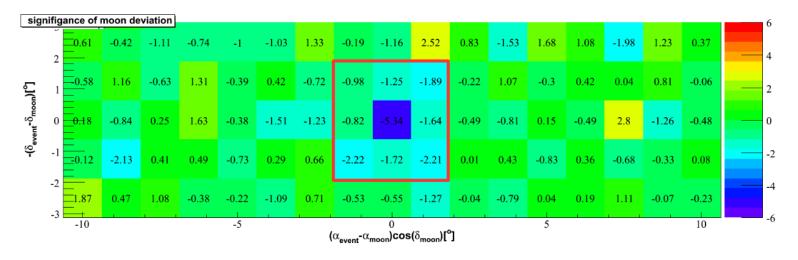
Levels of Data Challenge



 Data challenges at lower levels are more informative, but they are also more difficult

What We've Learned from the Moon

The moon blocks cosmic rays and acts as a muon calibration source

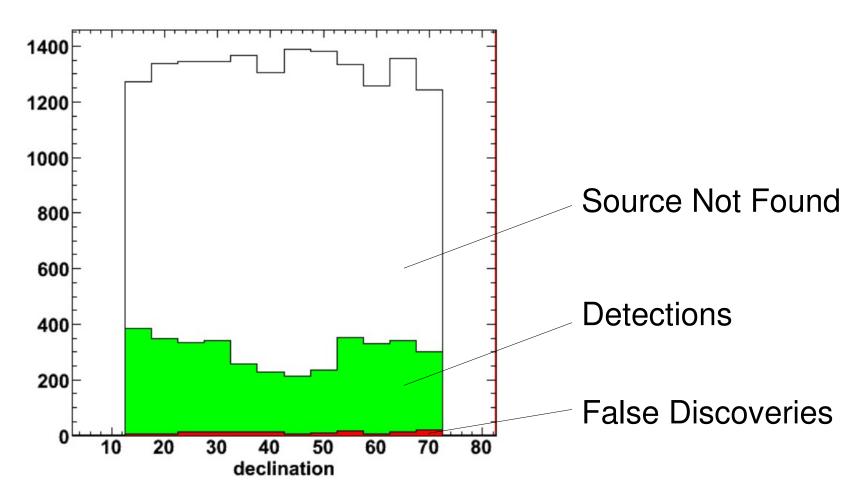


- We observe the moon shadow with a significance consistent with the expected value (5.34σ/6.1σ)
 - Evidence DAQ and Data Processing are working properly
 - Focus on analysis-level data challenges

- New neutrino point source method developed (FDR)
 - Advertised as powerful, but difficult to understand
- Need to assess the method in an unbiased way
 - Measure performance (Does the method work as advertised?)
 - If the method works, convince others
- Data challenge is an ideal strategy to accomplish this

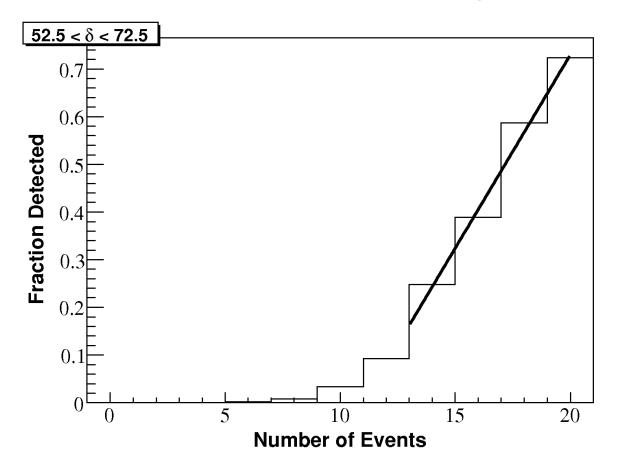
- Evaluate the method using a data challenge
 - Create ~17,000 sky maps with randomized AMANDA data and sources of various strength added
 - Send the set of maps to be analyzed
 - Receive coordinates for any sources discovered
 - Compare coordinates of sources with key: Sources listed within 5° of true location are considered detections; the rest are false discoveries.
- Software tools make this easy

FDR Data Challenge Results



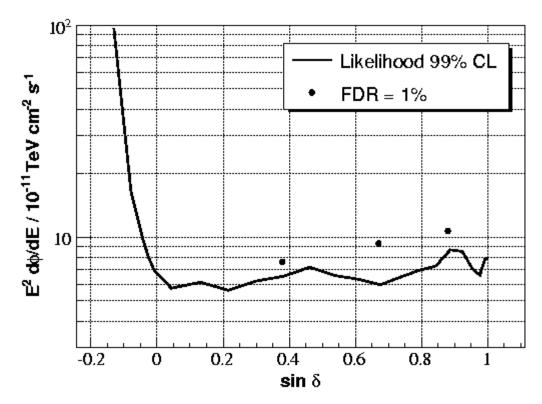
0.3% of sky maps contain false discoveries (Expected: < 1%)

Detection depends on simulated source strength



 Requires ~17 added signal events to have a 50% chance of detecting a source.

 Convert 50% discovery probability event thresholds to flux and compare against the baseline method



- FDR data challenge caught a problem:
 - Data had been assumed to be pure atmospheric neutrinos, but actually contains
 5% cosmic ray muon background
 - Tighter cuts remove the background and make the method consistent

Future Plans

- Use flasher events as a cascade data challenge
- Further analysis of the moon shadow
 - Systematic pointing errors
 - Measure muon point spread function
- Strongly encourage data challenges for most analyses when feasible

Summary

 IceCube is committed to using data challenges to ensure understanding of analyses and physics results

- Observation of the moon shadow has provided an end-to-end system test for muons
- FDR data challenge is a successful example of an analysis-level data challenge
- Flasher events will provide an end-to-end test for cascades